



Satellite 2012

Walter E. Washington Convention Center
Washington DC
March 12-14, 2012

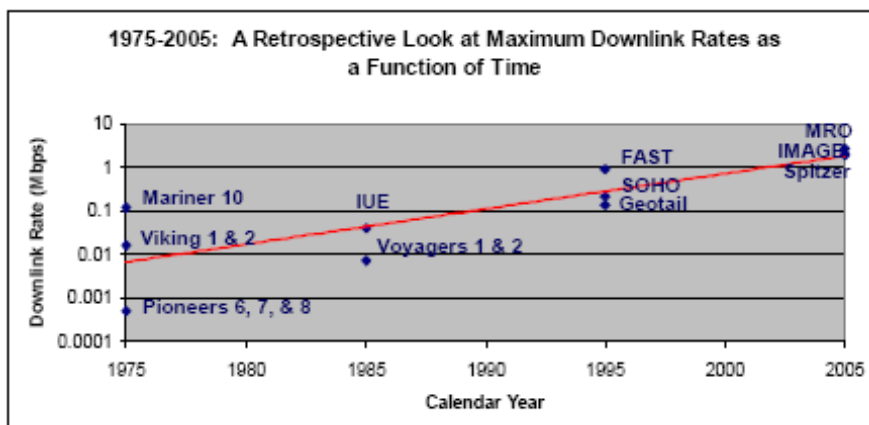
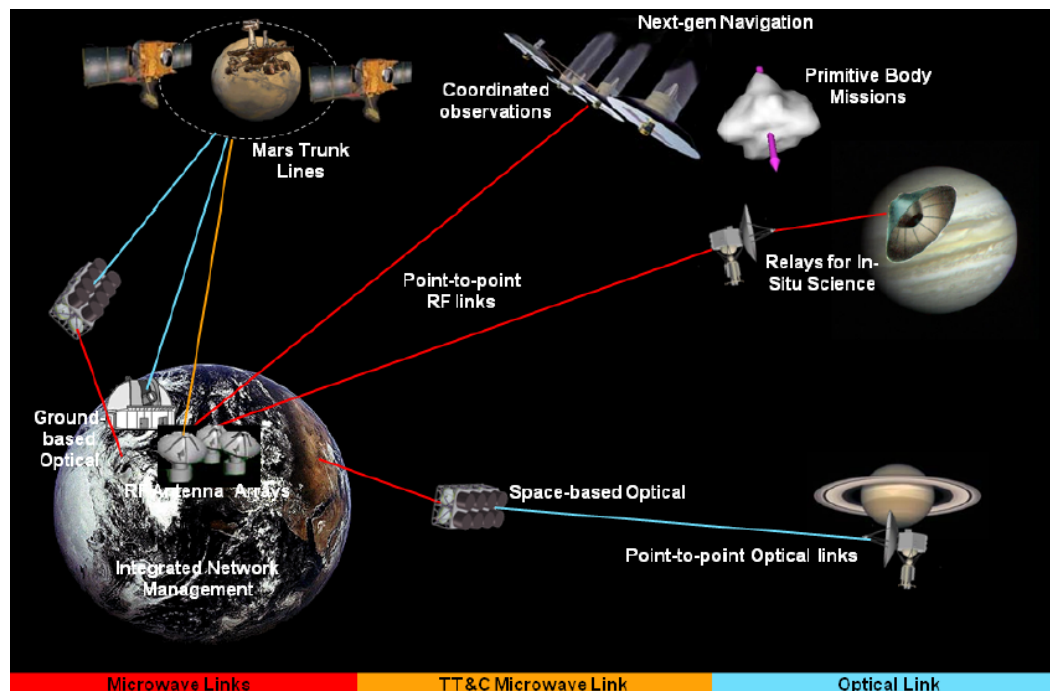
Technology Forum
New Methods of Antenna Design

Antenna Technology at NASA Glenn
Robert Romanofsky
robert.r.romanofsky@nasa.gov
NASA Glenn Research Center
Antenna and Optical Systems Branch

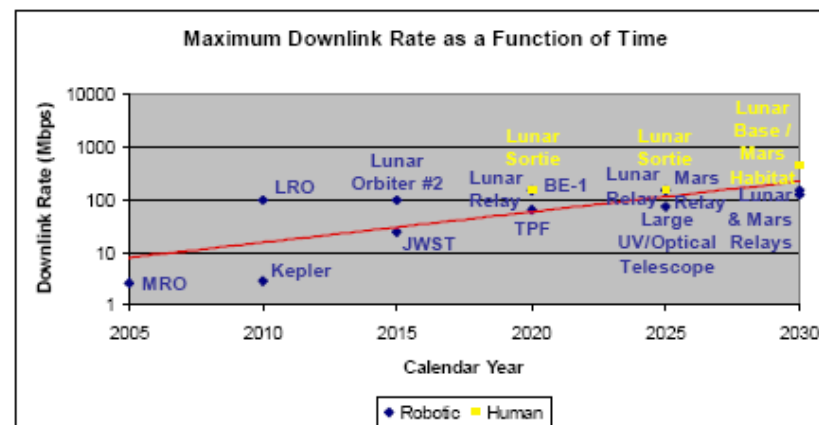


<http://aos.grc.nasa.gov/>

Space Communications and Navigation Project



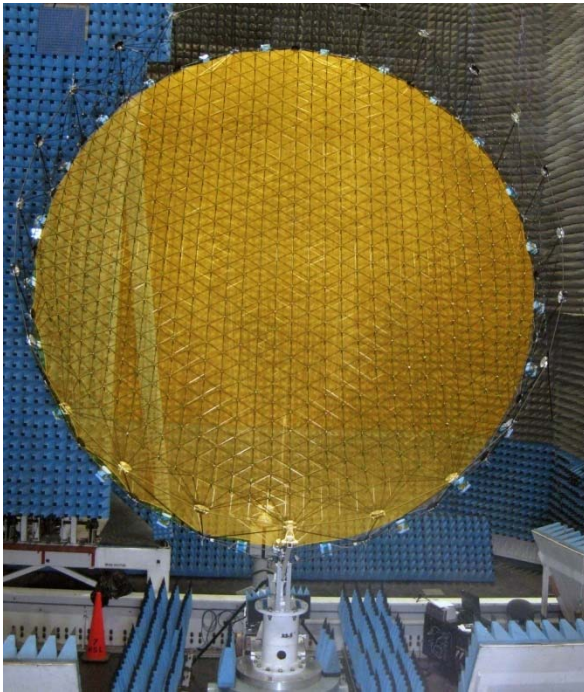
Deep-space communications downlink data rate trend



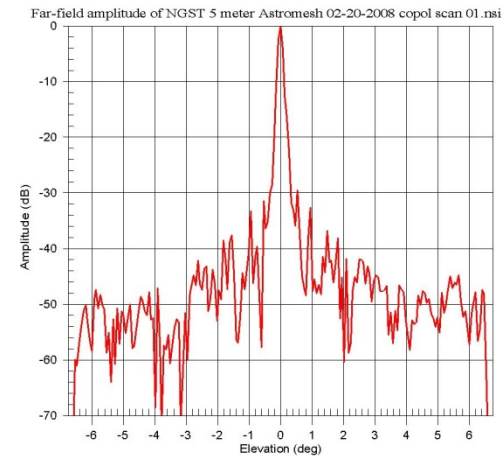
Expected Deep-space communications downlink data rate trend



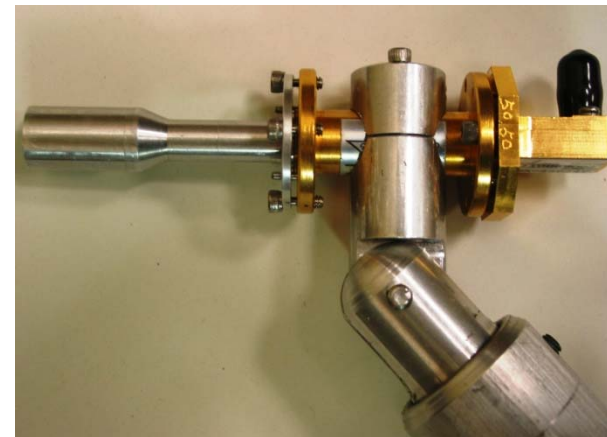
NGST¹ 5m Reflector Evaluated at 32, 38 and 49 GHz as well as laser radar surface accuracy mapping



NGST 5 m Reflector in NASA GRC Near-Field Range



Far Field Elevation pattern at 33 GHz
(Directivity = 62.8 dB)

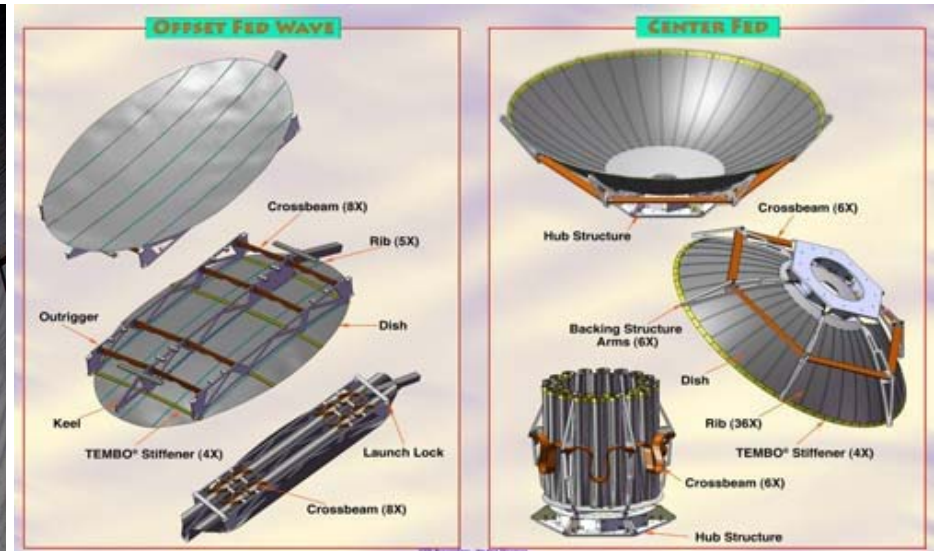
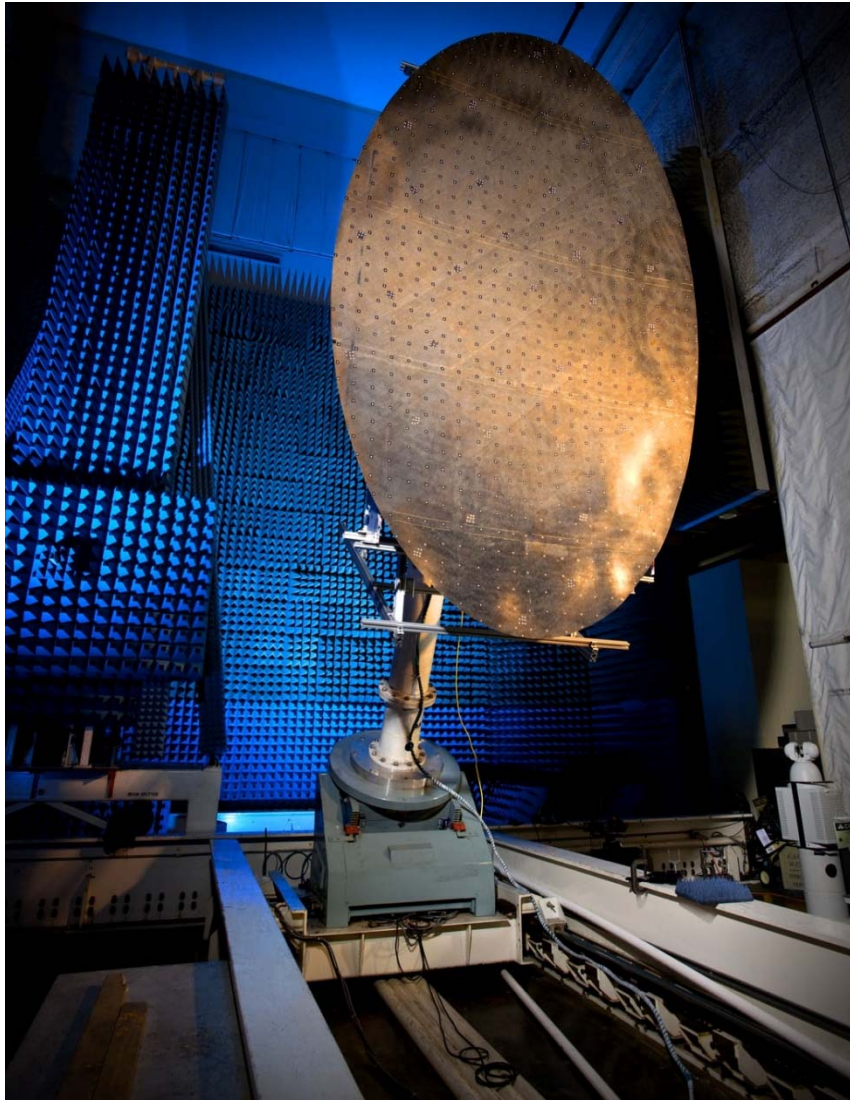


GRC Dual-band feed horn assembly

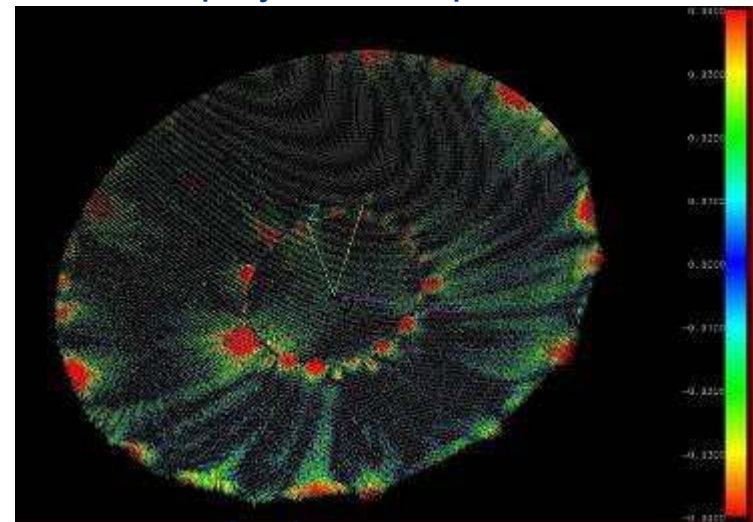


¹www.st.northropgrumman.com/astro-aerospace

Shape Memory Polymer Deployable Antenna



Deployment Sequence



4 m X 2.5 m Composite Technology Development Reflector In NASA Glenn Near Field Antenna Range

Laser Radar Surface Metrology on 3 m CTD Reflector at NASA Glenn

Self-Deploying Gossamer Support Structure

Phase II SBIR NASA Contract NNC06CA10C

Cornerstone Research Group, Dayton, OH



**Activation sequence of early prototype self-deploying truss:
one minute duration from compressed (2.5 in.) to deployed
(28.5 in.)**

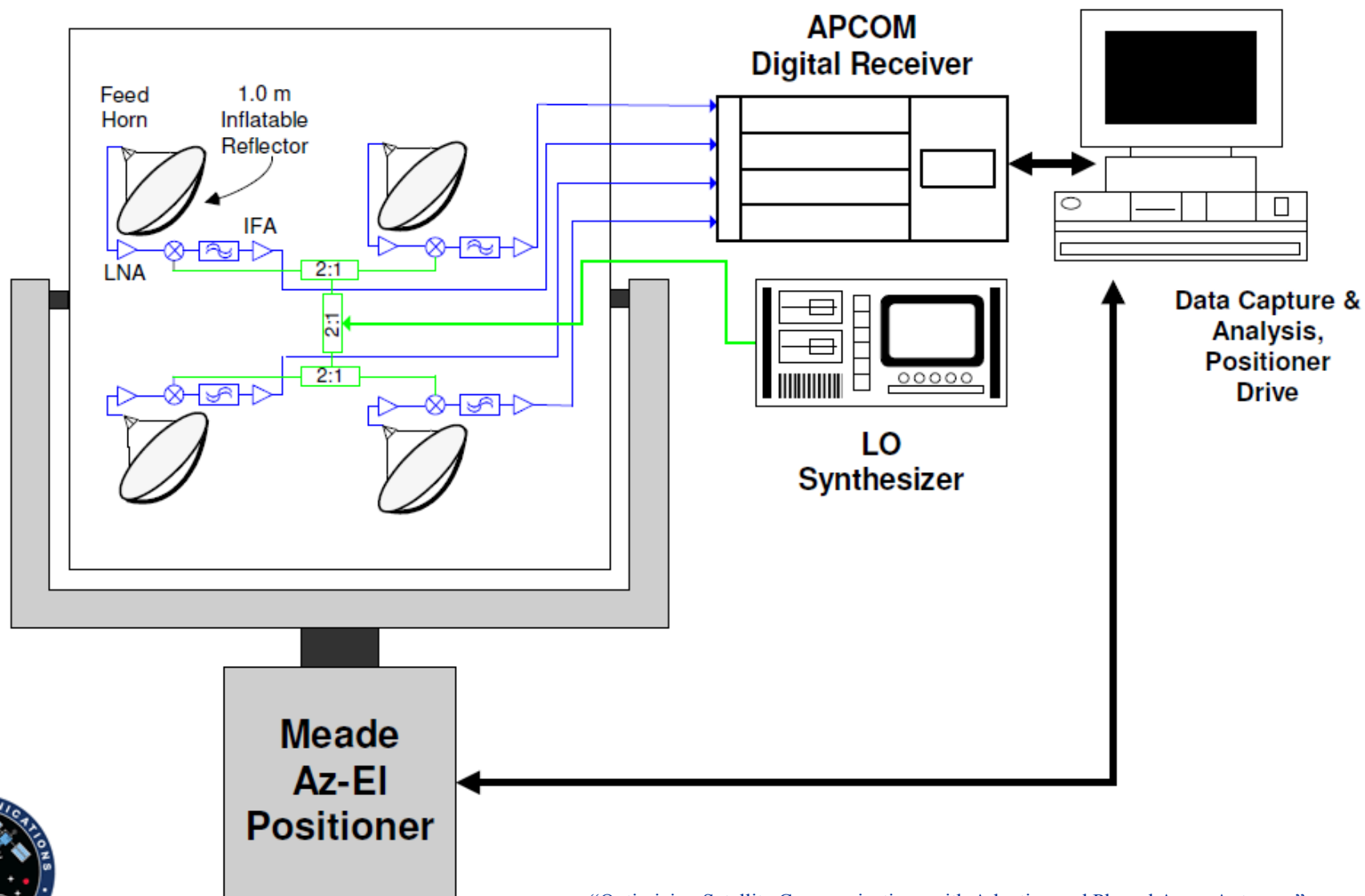


SAC-C Experiment





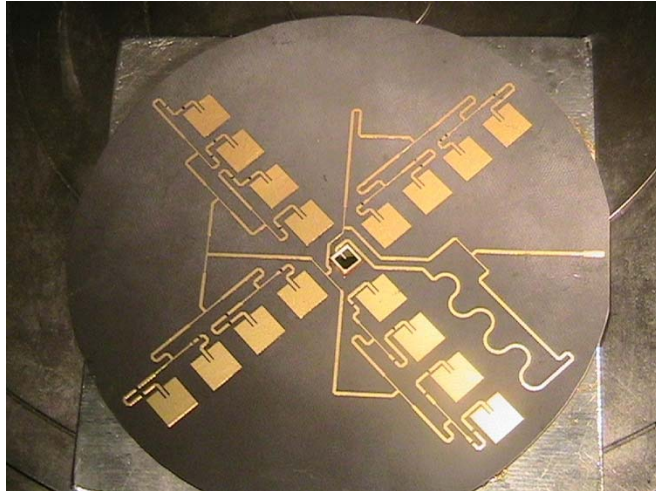
SAC-C Experiment¹



"Optimizing Satellite Communications with Adaptive and Phased Array Antennas",
M. Ingram, R. Romanofsky, R. Lee, et al., ESTC Conference, Anaheim, CA, June, 2004



Investigation of Noise Reduction Properties of the Zone Plate Antenna¹



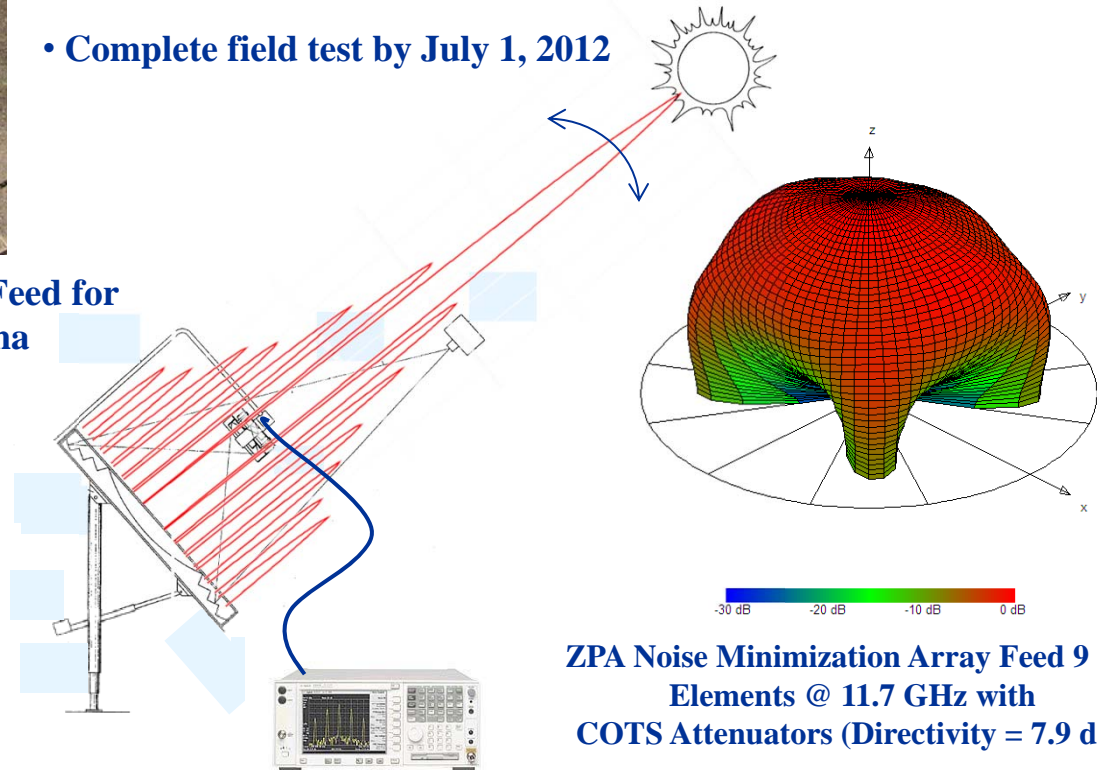
12 cm Uniform Amplitude Taper Array Feed for 150 cm diameter Zone Plate Antenna



Array Feed Mounted on ZPA Bridge

Autocorrelation of white noise at the output of the antenna is a sinc function. Thus for a band-limited system of 400 MHz, for example, the correlation time is approx. 2.5 ns. Since the ZPA induces a delay of 0.27 ns per ring by design, partial correlation of white noise may be present, resulting in some destructive interference of the noise signal

- Complete field test by July 1, 2012



Conduct solar flux and cold sky measurements and compare with well-characterized results to demonstrate noise cancellation of the ZPA.

¹alienworksltd@earthlink.net



SAA with GATR¹ Technologies since 2006 to develop Inflatable Radome Antenna System



- First-response disaster relief support
 - ✓ *humanitarian aid through support for communications in Haiti following the 2010 earthquake*
- Emergency search-and-rescue support
 - ✓ *The inflatable antenna was used to assist with communications efforts in the search for a missing girl in San Diego, California.*
- Large sports and entertainment venue backup communications
- Homeland security and military communications support
- Federal Laboratory Consortium Award for Excellence in Technology Transfer (2010)
- IR&D 100 Awards (2010)
- Featured in “Spinoff Day on the Hill” (2010)



¹www.gatr.com

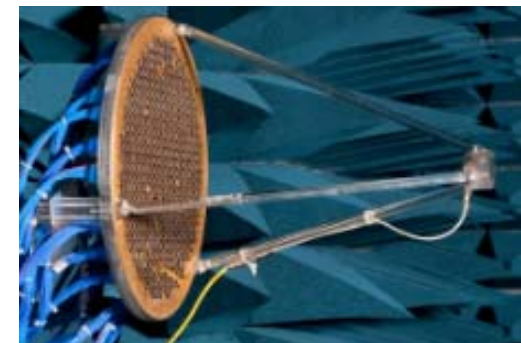
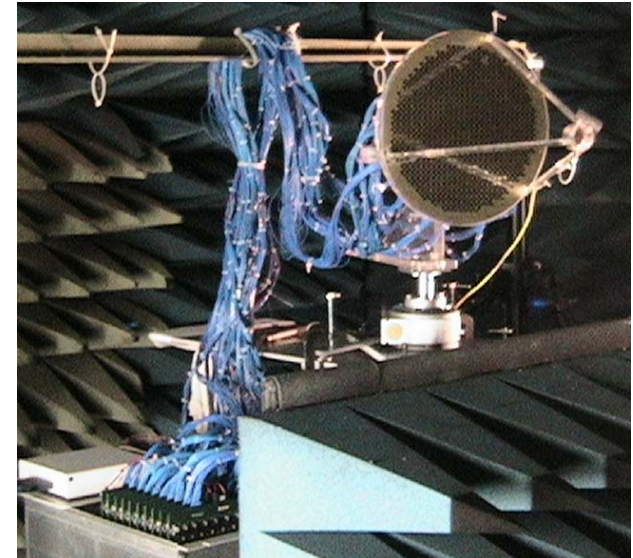


Officers with the Air Force Special Operations Command set up the inflatable antenna at Hurlburt Field in Eglin, Florida.



Ferroelectric Reflectarray

- A scanning reflectarray that consists of a flat surface with integrated phase shifters and patch radiators illuminated at a virtual focus
 - The signal passes through the phase shifters and is re-radiated as a focused beam in the preferred / target direction
 - Enabled by unique low loss ferroelectric phase shifters
- Reciprocal surface – the same aperture can transmit or receive so antenna is ideal for two-way communications or monostatic radar
- 2010 R&D 100 Award-winning innovation

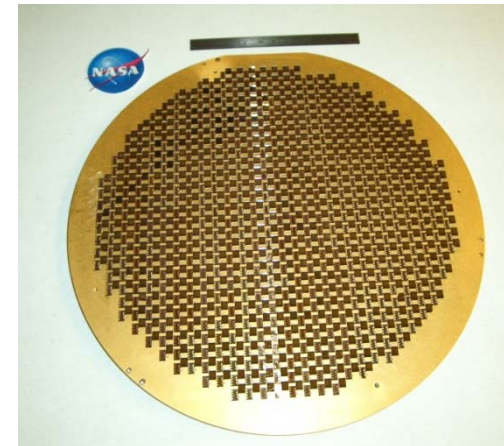




Ferroelectric Reflectarray

Key benefits are ...

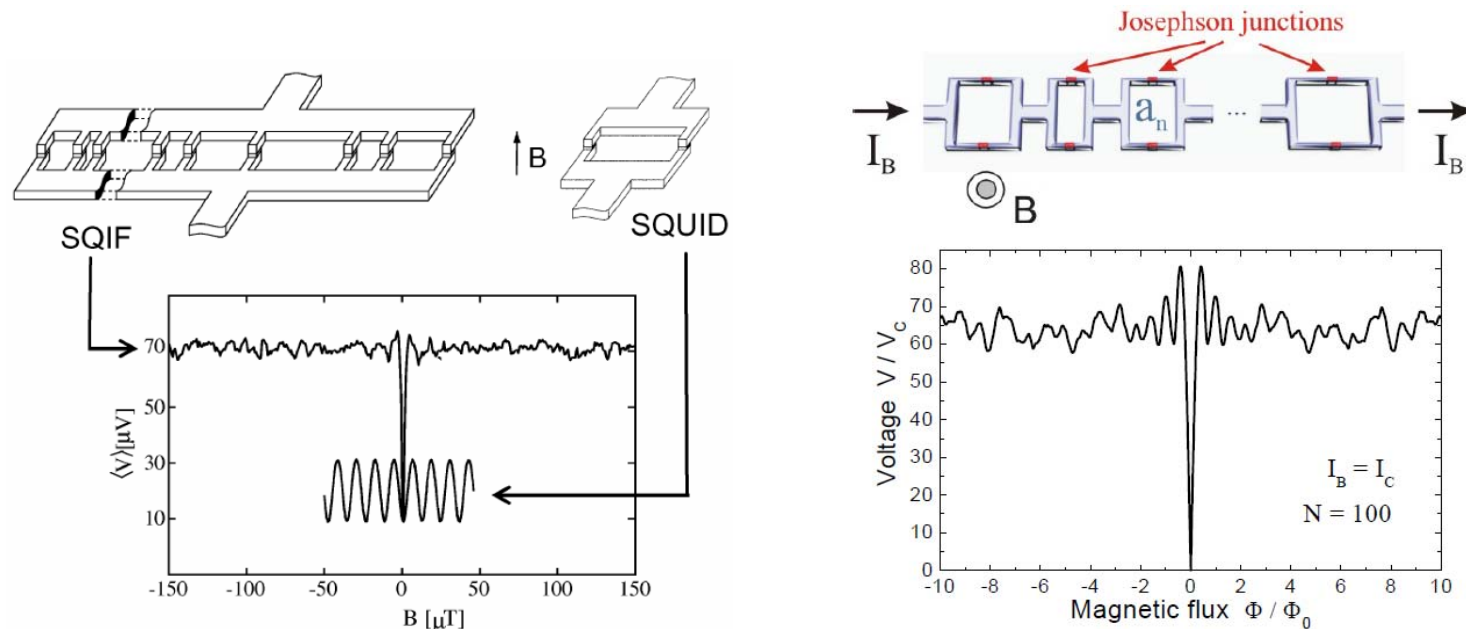
- **Significant cost reduction:** 10X to 100X (at scale) lower cost than direct radiating phased array
- **Unlimited gain:** gain is not limited by a beam forming manifold
- **Efficient:** up to a 5X reduction in power vs. GaAs MMIC array
- **Reliable:** high reliability compared to a gimbaled reflector due to no moving parts
- **Simple construction technique:** only three layers per phase shifter compared to at least five for an MMIC
- **Simple device lithography:** smallest feature size is $\sim 10 \mu\text{m}$ compared to sub-micron for MMIC devices



$\sim 28\text{cm}$ Active Diameter



Superconducting Quantum Interference Filter Receiver¹



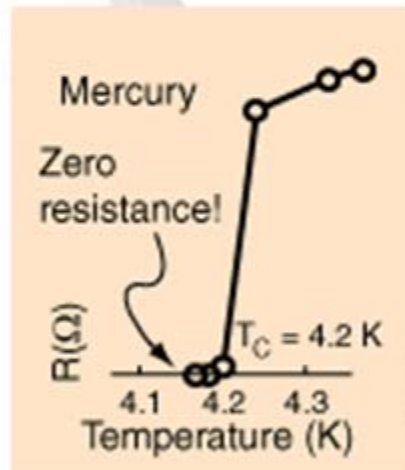
- Comparison between transfer characteristic of conventional SQUID (single loop) and SQIF (multiple-loop device). The amplitude of the peak is a function of the number (N) of SQUIDs in the array.

Sensitivity is a function of the slope and magnitude of the impulse-like response and dynamic range is a function of the linearity.



Table I. Critical Temperature for some key superconductors

Superconducting Material	Critical Temperature, T_c , (K)
Niobium (Nb)	9.2
Magnesium Diboride (MgB_2)	39
Yttrium Barium Copper Oxide ($\text{YBa}_2\text{Cu}_3\text{O}_7$)	90



Resistance versus Temperature for Hg¹ (perfect dc conductivity below a critical temperature)

- The microscopic origin of this pure *quantum mechanical* phenomenon is the formation of pairs of two conduction electrons having opposite momentum and spin (Cooper pairs)
- In conventional superconductors, electron attraction is generally attributed to an electron-lattice interaction. In the BCS framework, superconductivity is a macroscopic effect which results from "condensation" of Cooper pairs. The attractive interaction between electrons (necessary for pairing) is brought about indirectly by the interaction between the electrons and the vibrating crystal lattice (phonons)
- The spatial extent of the pair correlation can be several orders of magnitude larger than the interatomic distance and is characterized by the coherence length ξ ,
- Due to the quasi-boson character of the Cooper pairs, they are allowed to condensate into a common ground state, which can be described by a single macroscopic wave function

¹<http://hyperphysics.phy-astr.gsu.edu/hbase/solids/scond.html>



Semiconductor Receiver versus SQIF Receiver

\mathcal{E} vs. \mathcal{B}

SQUIDS can detect magnetic fields lower than one flux quantum $h/(2e)(\approx 10^{-15} \text{ Wb})$, 10^{-18} Wb reported in the literature

Conventional Receiver

Mars link at 64 MBPS

- EIRP = 84 dBW ($\approx 2.5 \times 10^8 \text{ W}$)
Assumes 100 W TWT, 12 m aperture
Range = $3.7 \times 10^8 \text{ km}$
- Power density at receiver $\approx 2.8 \times 10^{-16} \text{ W/m}^2$
Electric Field $\approx 4.6 \times 10^{-7} \text{ V/m}$
Displacement flux density $\approx 10^{-18} \text{ C/m}^2$
- Receive Antenna Aperture
QPSK, Block Turbo Code, 3 dB margin
Required $E_b/N_0 = 4.6 \text{ dB}$

72 meters

SQIF Superconducting Receiver

Mars link at 64 MBPS

- EIRP = 84 dBW ($\approx 2.5 \times 10^8 \text{ W}$)
Assumes 100 W TWT, 12 m aperture
Range = $3.7 \times 10^8 \text{ km}$
- Power density at receiver $\approx 2.8 \times 10^{-16} \text{ W/m}^2$
Electric Field $\approx 4.6 \times 10^{-7} \text{ V/m}$
Displacement flux density $\approx 10^{-18} \text{ C/m}^2$
Magnetic Field $\approx 10^{-9} \text{ A/m}$
Magnetic flux density $\approx 10^{-15} \text{ Wb/m}^2$
- Receive Antenna Aperture
Flux Concentrator
Mechanical refrigerator at 4K

$\approx 1 \text{ meter???$